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Bar Spread Functions for Parabolic Image Motion

The effect of parabolic image motion in panoramic cameras becomes significant for large obstruction ratios.

INTRODUCTION

 S ingle bar analysis recently has acquired importance in the assessment of optical systems and is of direct importance in aerial photography1 because the projected silhouettes of the objects such as pipes, rails, white road lines, highways, etc., are effectively single bar objects and their imagery is of considerable importance. For an infinitesimally thin bar, one obtains the line spread function. On the other hand, with a very

thus, various aspects of the images formed by annular apertures need extensive investigation. An extensive list of relevant references is given elsewhere^{7,8}. Some recent relevant studies may also be noted⁹⁻¹⁵.

Most of the investigations have dealt with diffraction/aberration limited systems. However, greater and greater precision is being demanded of motion sensing and image motion compensation (IMC) equipment as precision in aerial reconnaissance,

ABSTRACT: A *graphical analysis has been made ofthe image intensity distribution for bar targets of varying widths in presence of parabolic image motion for various sizes of the pupil-obscuration in panoramic cameras. Asymmetry in the intensity distribution in the image has been investigated in detail. For large obstruction ratios, deleterious effects of motion become very significant and the tolerance on the image motion appears to become very stringent. Results also have been obtained for the image oftwo other isolated extended objects, namely a triangle and an object in the shape ofa single cycle of pure sinusoid. The line spread function has also been obtained and single bar contrast functions plotted.*

wide bar the image may be similar to that of an edge. A bar image can also be synthesized as the difference of the two edge spread functions¹. A single bar also has found application as a resolving power test target¹³ and the methods of defining and calculating the contrast of bar images have been developed 4-6. Generally, the imaging system is considered to have a circular aperture.

Annular aperture systems are also of considerable importance and more interesting than the conventional circular aperture systems. The central portion of the Airy pattern in such a system becomes narrower and thus increases the depth of focus of the system;

surveillance and other intelligence activities increases. Since IMC techniques are never perfect, some residual image motion¹⁶⁻²⁰ is always present and for a significant class of problems puts a limit to the resolution. Vibrations from machinery close by, the spinning or tumbling of the spacecraft, the jerking of the camera by the hand of an astronaut and even the click of the camera shutter are some typical sources of image motion¹⁶ through 20. Effects of linear and sinusoidal motion on the images of various objects have been extensively investigated $21 - 23$ in the past.

Interest in transverse scanning panoramic aerial cameras¹⁷⁻²⁴ has stimulated con-

siderable analysis of their performance characteristics. Two forms of transverse scanning panoramic cameras are available. The first has film fixed upon a cylindrical focal surface and sweeps the image across the film by swinging the lens itself; the second has a fixed lens and slit arrangement accomplishing the transverse scan by rotating a mirror or a prism in front of the lens and synchronising this with the motion of the film past the slit. The image motion in these is compensated by moving the optics or the film in some way that will minimize the over-all image motion.

Residual image motion in transverse scanning panoramic cameras results from the finite slit width which precludes the film's IMC from being exactly correct for all images simultaneously exposed in the slit. The image displacement is a parabolic function of time¹⁷⁻²⁶. Such motion recently has been considered with respect to its effect on imagery²⁶ holographic measurements²⁷ and image restoration^{28,29}.

We have discussed the single bar targets in the presence of parabolic image motion for annular aperture systems. The results for a triangle and an object having the shape of a single cycle of pure sinusoid are also included.

THEORY

The image intensity spectrum $I(\omega)$ and the object intensity spectrum $O(\omega)$ are related by

$$
I(\omega) = T(\omega) O(\omega) \tag{1}
$$

where $T(\omega)$ is the optical transfer function of the system and *w* the normalised spatial frequency. Intensity distribution *i(V)* in the image is the inverse Fourier transform of $I(\omega)$. Hence

$$
i(V) = \int_{-\infty}^{\infty} T(\omega) O(\omega) \exp(iV\omega) d\omega \qquad (2)
$$

where V is the distance in the image plane in diffraction units*. Since a one-dimensional Fourier transform has been written in Equation 2, the formula is applicable only to images of one-dimensional objects. The object spectrum $O(\omega)$ can be obtained by taking the Fourier transform of the object intensity function $O(V)$. For a single bar^{22,30} this turns out to be

$$
O(\omega) = (\pi \omega)^{-1} \sin (\omega L) \tag{3}
$$

* $V = \pi x / \lambda F$ where *x* is linear distance in the image plane from optical axis, λ is wavelength of radiations, and F is the f-number of the optical system under consideration.

where $2L$ is the width of the bar in diffraction units. Since $T = T_r - iT_i$ where T_r and T_i are respectively the even and odd function of ω , $i(V)$ reduces to

$$
i(V) = 2(\pi)^{-1} \left[\int_{0}^{2} T_r(\omega) \sin (\omega L) \cos (V \omega) \right]
$$

$$
(\omega^{-1}) d\omega + \int_{0}^{2} T_i(\omega) \sin (\omega L) \sin (\omega l)
$$

$$
(V \omega) (\omega^{-1}) d\omega
$$
 (4)

Since $T(\omega)$ vanishes for $\omega > 2$, the limits of the integral have been taken from 0 to 2. The optical transfer function of a diffraction limited system may now be multiplied by the optical transfer function of the motion to give the transfer function of a system afflicted with motion-induced image degradation. Hence

$$
T(\omega) = D(\omega) C(\omega) \tag{5}
$$

where $D(\omega)$ is the diffraction limited MTF and $C(\omega)$, the motion OTF.

The expression for $D(\omega)$ in case of an annular aperture is well known^{7,8}.

Further, the OTF for parabolic motion^{17,24,25} is given by

$$
C(\omega) = \frac{\sqrt{I_1^2 + I_2^2}}{2\sqrt{\omega A}} e^{-i \theta'} \tag{6}
$$

Here $\theta' = \tan^{-1}(I_2/I_1)$ is the phase lag and I_1 and I_2 are the Fresnel's integrals given by

$$
I_1 = \int_{0}^{2\sqrt{\omega A}} \cos(\pi/2 u^2) du
$$

$$
I_2 = \int_{0}^{2\sqrt{\omega A}} \sin(\pi/2u^2) du
$$
 (7)

A is the dimensionless motion parameter given by $A = Bte^2/2\lambda F$. Here B is defined by the equation of the parabolic motion¹⁷, *te* is total exposure time, λ is the wavelength of the incident light, and F is the f-number of the optical system.

Therefore the final expression for intensity distribution for a single bar in the presence of motion becomes

$$
i(V) = 2(\pi)^{-1} \int_{0}^{2} D(\omega) \frac{\sqrt{I_1^2 + I_2^2}}{2\sqrt{\omega A}} \cos(V\omega - \theta')
$$

$$
\cdot \sin(\omega L) (\omega^{-1}) d\omega
$$
 (8)

The object spectrum for a triangular-

shaped object³⁰ and an object in the shape of a single cycle of a sinusoid are given respectively by

$$
O(\omega) = (\pi \omega^2)^{-1} \sin^2(\omega L) \tag{9}
$$

and

$$
O(\omega) = \left[2\omega\left(\pi^2 - 4\omega^2L^2\right)\right]^{-1} \pi \sin\left(2\omega L\right) (10)
$$

By using these relations respective images can be calculated.

However, there is a complication in defining and computing the modulation of the image because any peak value against a background of 0.0 gives a modulation of 1.0 when calculated according to the conventional formula, namely,

 $\text{Modulation} = (I_{max} - I_{min})/(I_{max} + I_{min})$

where *[max* is the peak value and *[min* is the background. Williams' has shown that the final modulation of a single bar target of original modulation Mo, which is degraded according to some MTF, is equl to $b \times Mi$ where *b* is the final peak value in normalised intensity units of an equal size bar image with an input intensity and modulation of 1.0, which has been degraded by the same MTF. Therefore, the final modulation of bar images which had an input intensity and modulation 1.0 is defined to be their maximum intensity or effective exposure value b. Considering this definition as an extension from the cases where the input modulation is less than 1.0, this definition is very clear. It must be noted that this definition may be extended to define a logarithmic contrast as was done by Charman⁵. However, we characterize our images by the ordinary contrast, i.e., the modulation.

RESULTS AND DISCUSSIONS

Intensity distributions were calculated numerically on an electronic computer using 32-point Gauss Quadrature method for $2L =$ 2.0, 4.0, 8.0, 12.0, and 16.0. Some representative results of calculations are shown graphically in Figures 1 through 4 for bar widths 2L = 2.0 and 8.0 diffraction units. The results in the absence of image motion $(A = 0.0)$ are also shown for the sake of comparison by the dotted curves in Figures 1,3, and 4. Figure 1 demonstrates the influence of parabolic image motion on the bar images formed by annular aperture systems with an obstruction ratio $\eta = 0.25, 0.50,$ and 0.75.

In all the three cases, it is observed that the secondary maxima in the motion-free images arising due to diffraction at the annular aperture are no longer visible in the

FIG. 1. (a) Images of bar in the presence of parabolic motion for L = 1.0, $\eta = 0.25$ for A = 0.0, 0.5, 1.0, 1.5, and 2.0.

FIG. 1 (b) Same as 1(a) for $\eta = 0.50$.

motion-degraded images and degradation becomes more deleterious for higher values of η ; also, there is broadening of the diffraction figure with increasing value of η and A. It must therefore be stressed that increasing obscuration leads to more stringent tolerance on the acceptable level of motion. This is because the contrast in the images of extended objects formed by systems with pupil obscuration is reduced cumulatively by the obscuration and the image motion. Clearly, more efficient image-motion compensation/ stabilization techniques must be developed for annular aperture systems than are availa-

FIG. 1 (c) Same as 1(a) for $\eta = 0.75$.

FIG. 1 (d) Same as 1(a) for L = 4.0 and $\eta = 0.25$. FIG. 1 (f) Same as 1(a) for L = 4.0 and $\eta = 0.75$.

FIG. 1 (e) Same as 1(a) for $L = 4.0$ and $\eta = 0.50$.

ble for conventional circular aperture systems.

It is interesting to note from Figure 1 that a shift in the position of peak intensity is observed in going from $A = 0.0$ to 0.50 but it remains constant for higher values ofA. This is explained on the basis of the fact that the phase part of the transfer function of parabolic motion becomes oscillatory beyond a certain value of ω for given value of A (for example at $\omega = 0.7$, for $A = 1.0$) and then becomes constant afterwards in the high frequency region. Intensity distribution is asymmetric with respect to origin in the presence of motion because θ' is a nonlinear function of w. Such asymmetry is also observed in investigating the comatic images by Ingelstam *et al.*³¹ and Barakat³².

Figure 1 also shows that the central intensity increases as the width of the bar is increased for the same amount of motion. Intensity decreases with the increasing value of motion parameter for the same bar width. For large values of L the bar object behaves as disk or an edge object. Figure 2 shows the plots of the modulation of the final images as a function of the image-motion parameter for different values of L when the object is in the form of a single cycle of pure sinusoid, a triangle, and a bar. It can easily be seen that the contrast is greater for a bar target than for a sinusoidal or triangular targets of the same

FIG. 2 (a) Contrast in the images of bar, sinusoidal and triangular object with L = 1.0 for $\eta = 0.25, 0.50$ and 0.75.

period. This is due to the fact that square wave shape carries more energy than the narrow-peaked sine wave at the same period, and hence degrades less when both are blurred by convolution with the same spread function. However, the contrast becomes nearly the same for smaller values of L and η ; therefore the difference in the images of these types of isolated extended but aperiodic objects are not considerable for low values of L and η . The difference is, however, appreciable for higher values.

Figure 3 shows the plot of intensity distribution for very narrow bar $(L = 0.1)$ and

F_{IG}. 2(b) Same as $2(a)$ for $L = 4.0$.

FIG. 3 (a) Image of a line with $L = 0.1$, $\eta = 0.25$ for $A = 0.0, 0.5, 1.0, 1.5,$ and 2.0.

FIG. 3 (b) Same as 3(a) for $\eta = 0.50$.

gives the line spread function for annular aperture system.

In Figure 4 contrast is plotted as a function of normalised frequency for various values of A and η . We find that the contrast never falls to zero for a bar of any finite width; however, it reduces with increasing values of η and A. The wide spectral bandwidth of a single bar provides an explanation for the frequent appearance in aerial photography of isolated objects.

CONCLUDING REMARKS

Rosenau,²⁴ while considering the perfor-

FIG. 3 (c) Same as 3(a) for $\eta = 0.75$.

FIG. 4 (a) Single bar contrast transfer function for $\eta = 0.25$ for A = 0.0, 0.5, 1.0, 1.5, and 2.0.

mance of a model 501 light weight aerial panoramic camera, derived an expression for co-efficient *B* in the form

$$
B = \frac{1}{2} (v/h) \dot{x}_f \sin \beta
$$

where *v* is aircraft velocity relative to earth, h is altitude of film plane above the earth, \dot{x}_f is the film velocity to compensate image motion, and β is scan angle. In a typical case, the camera was exposed for 1/55 sec at a maximum rate of 0.30 rad/sec with a film transport velocity of lOA in/sec. Thus, for a 60° scan angle, the total image movement during the time of exposure becomes

FIG. 4 (b) Same as 4(a) for $\eta = 0.50$.

FIG. 4 (c) Same as 4(a) for $\eta = 0.75$.

B $te^2 = (1/88)$ mm

Since in most of the photo-reconnaissance missions the sensor is kept at a very high altitude (as in a U-2 type aircraft or a satellite), a long focal lens having a large F-value is required¹⁷. Thus for λ = 5000A° and $F = 5$ to 10 (say), the motion parameter

$A = Bte^2/2\lambda F = 2.28$ to 1.14

This shows that the values of A considered in our computations are realistic.

It is now well known that the transfer

function of certain optical systems with appropriate amplitude filters is very similar to the transfer function of annular aperture systems³³. Hence, the trends for motiondegraded bar spread function in these will be similar to our results. It may also be helpful in various applications using reflecting components such as in space optics 34 , night vision aids³⁵, low-light-level systems³⁶, optical processing¹⁷ of images, etc.

We would also like to mention that the motion-degraded images oftwo-dimensional objects such as disks, annuli, etc., cannot be calculated by using Equation 5 because the OTF of the total system does not possess a rotational symmetry but acquires a certain azimuthal dependence. Work in this context has been done by Rattan and Singh³⁷ and Gupta and Singh³⁸.

The problem of restoring images degraded by parabolic motion is under investigation and results will be reported at a later date.

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